



Institute for Advanced Study
Technical University of Munich



Theseus SXI and AGN

Paolo Giommi

Active Galactic Nuclei: what's in a name?

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B. De Marco · P. Giommi · R. C. Hickox · G.
T. Richards · V. Smolčić · E. Hatziminaoglou ·
V. Mainieri · M. Salvato

The Astronomy and Astrophysics Review
2017 arXiv 1707.07134

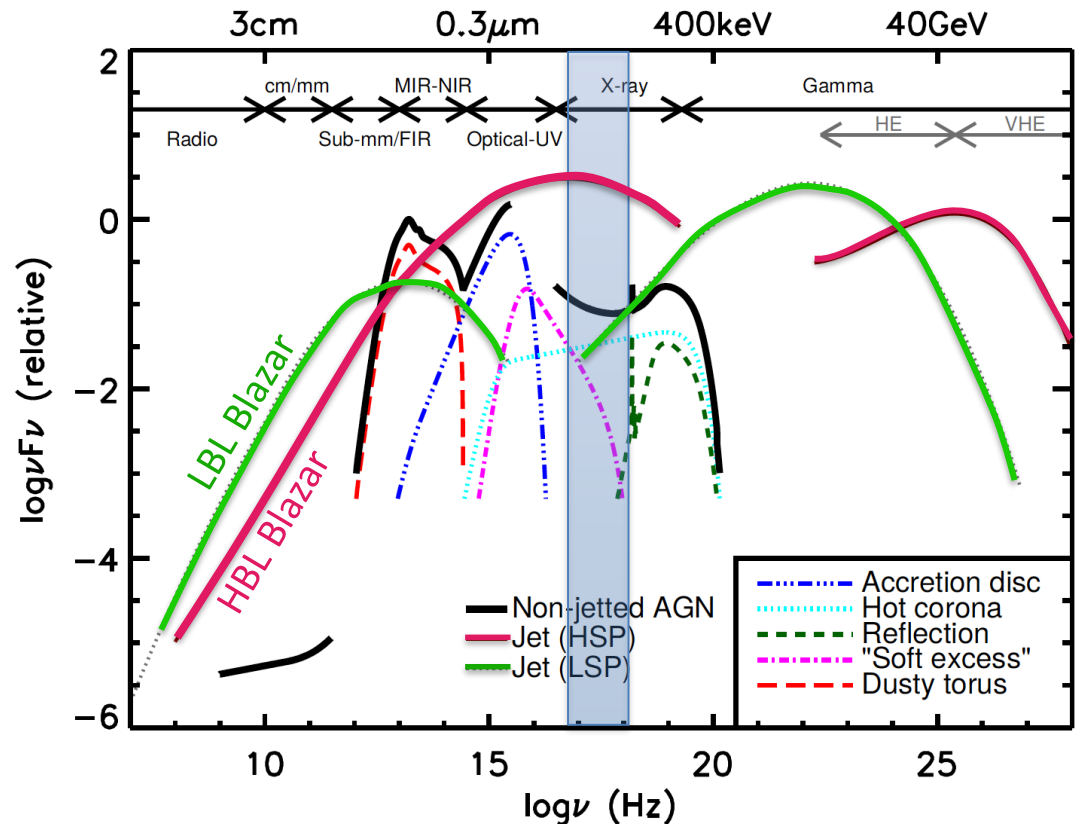
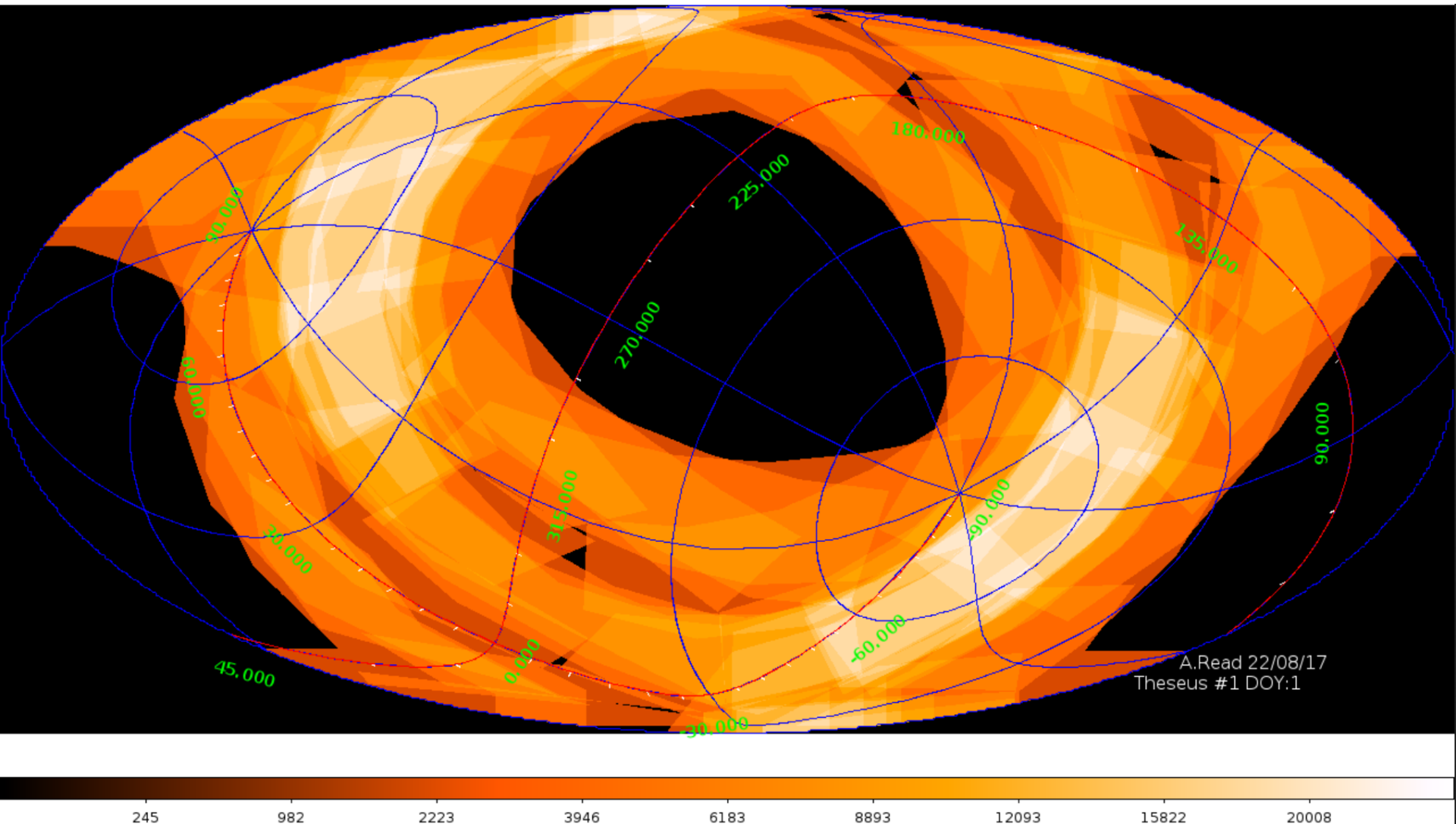
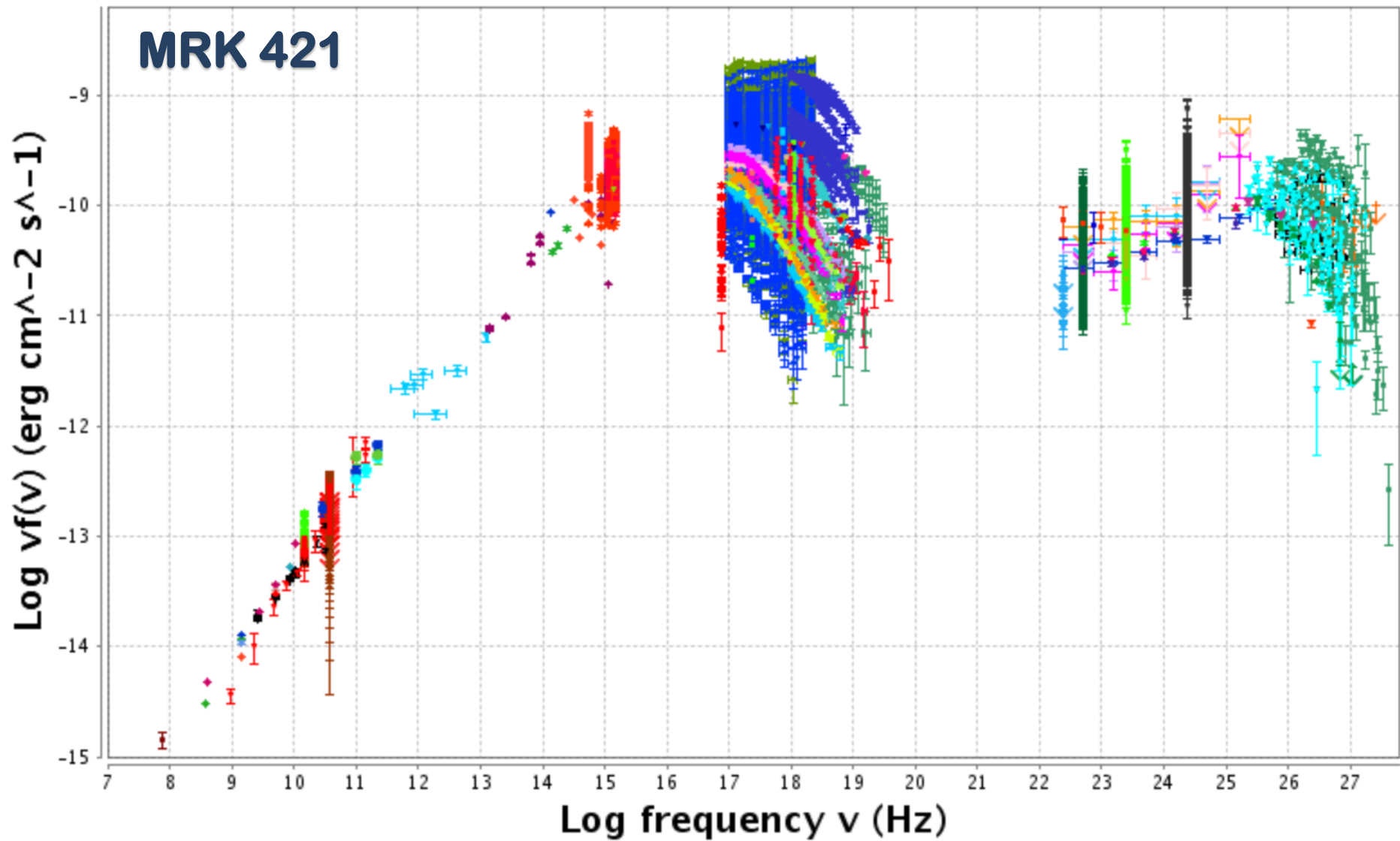


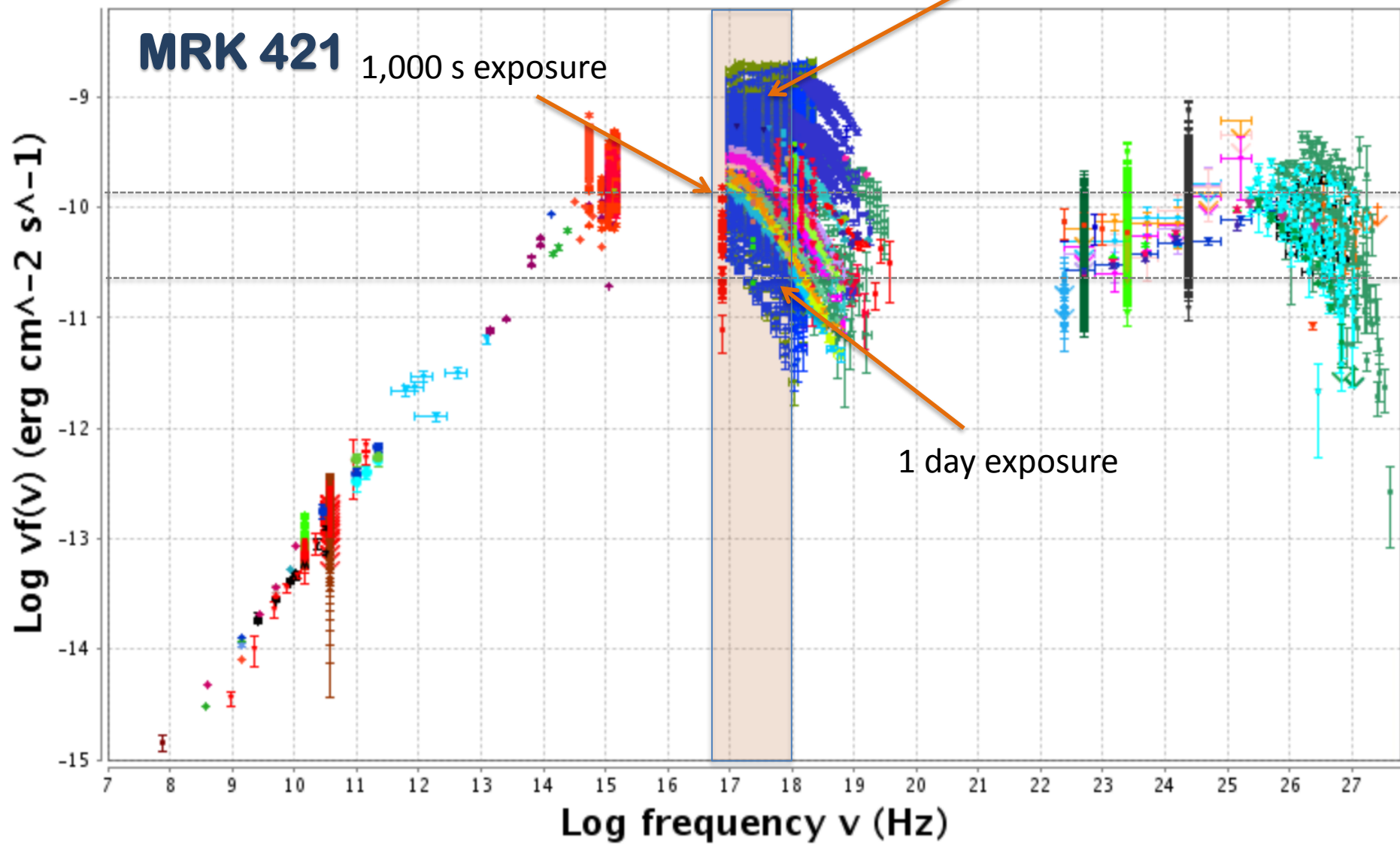
Fig. 1 A schematic representation of an AGN spectral energy distribution (SED), loosely based on the observed SEDs of non-jetted quasars (e.g. Elvis et al., 1994; Richards et al., 2006a). The black solid

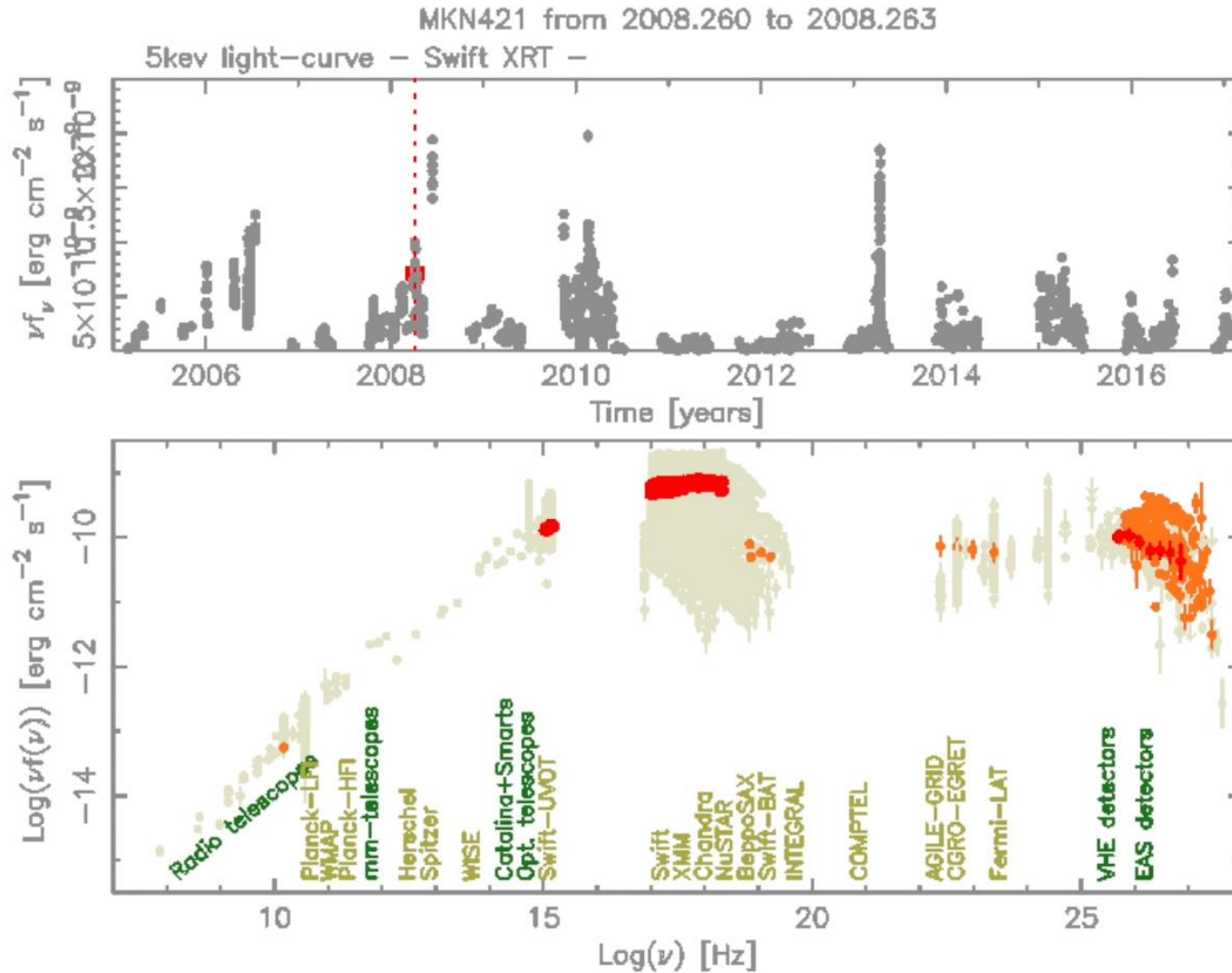


SXI 1 day exposure











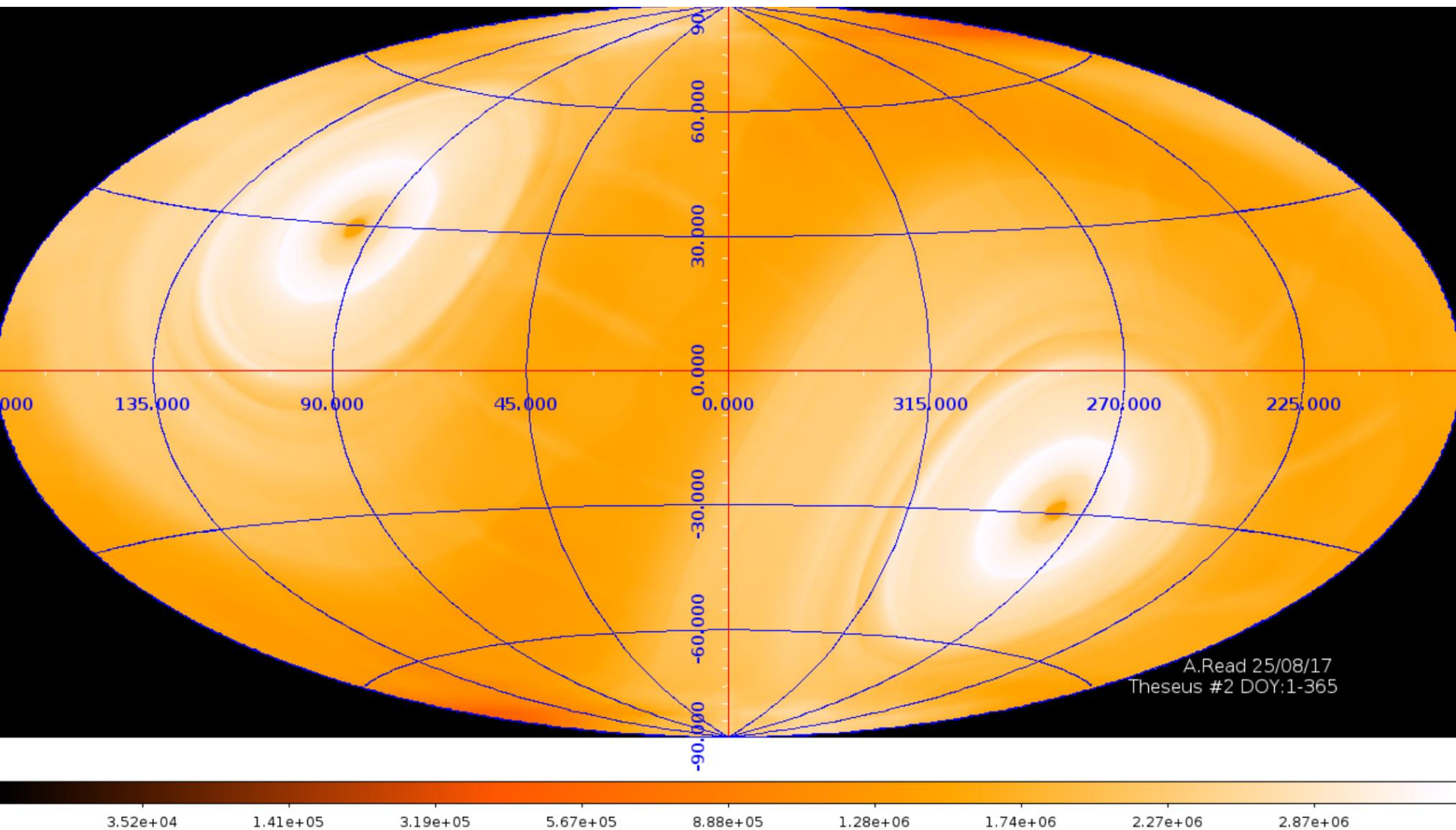
**The brightest source in the
γ-ray sky (for a few days)**

The blazar 3C454.3



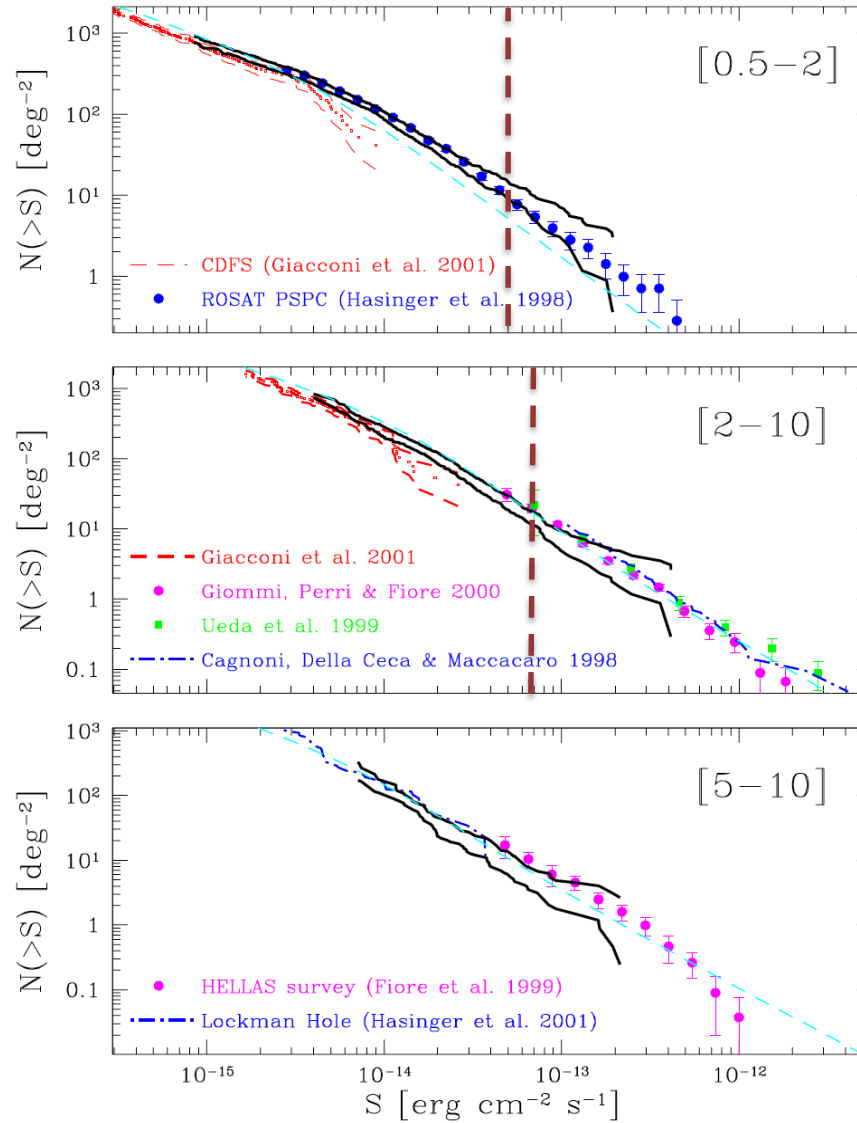


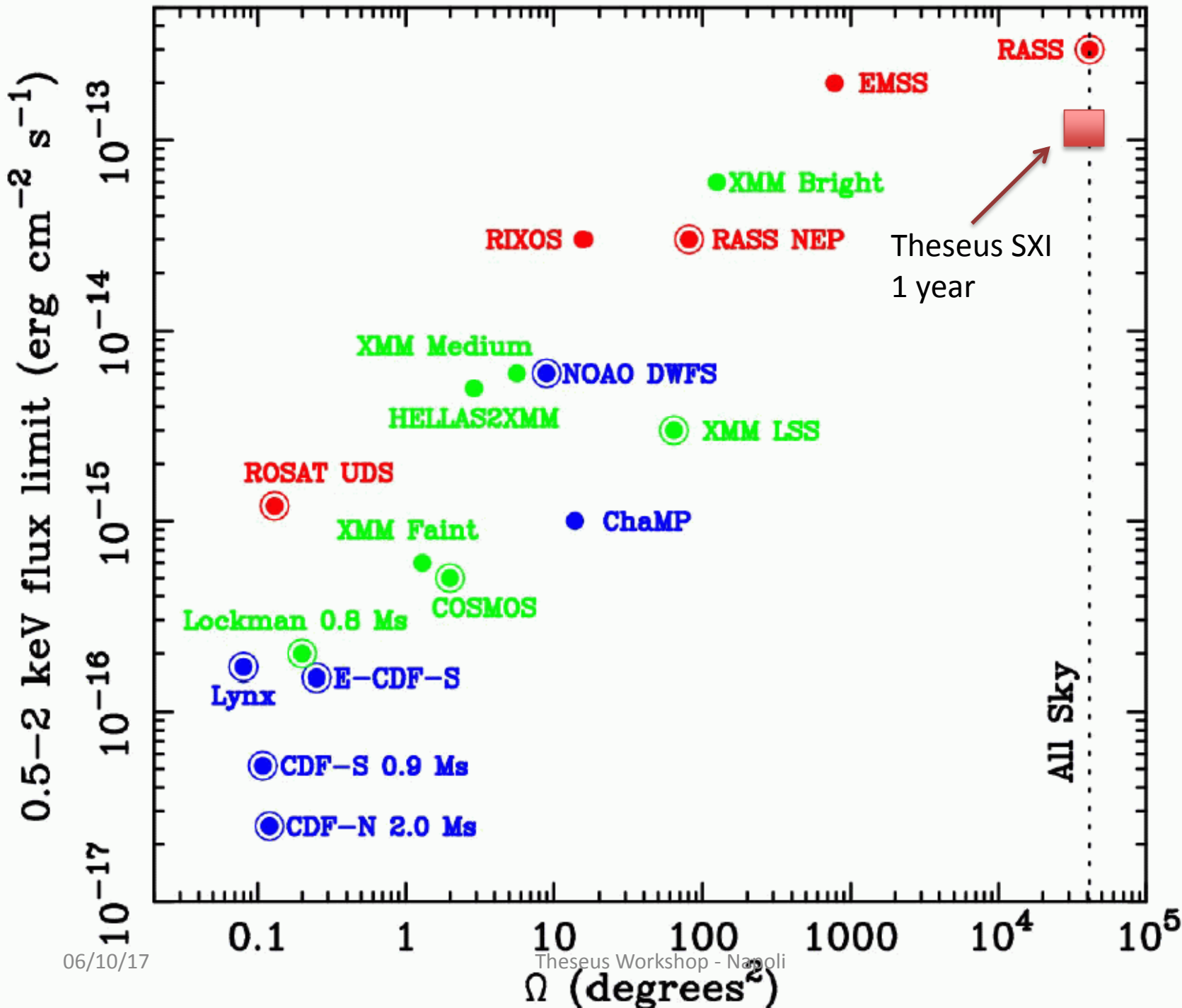
SXI 1 year exposure

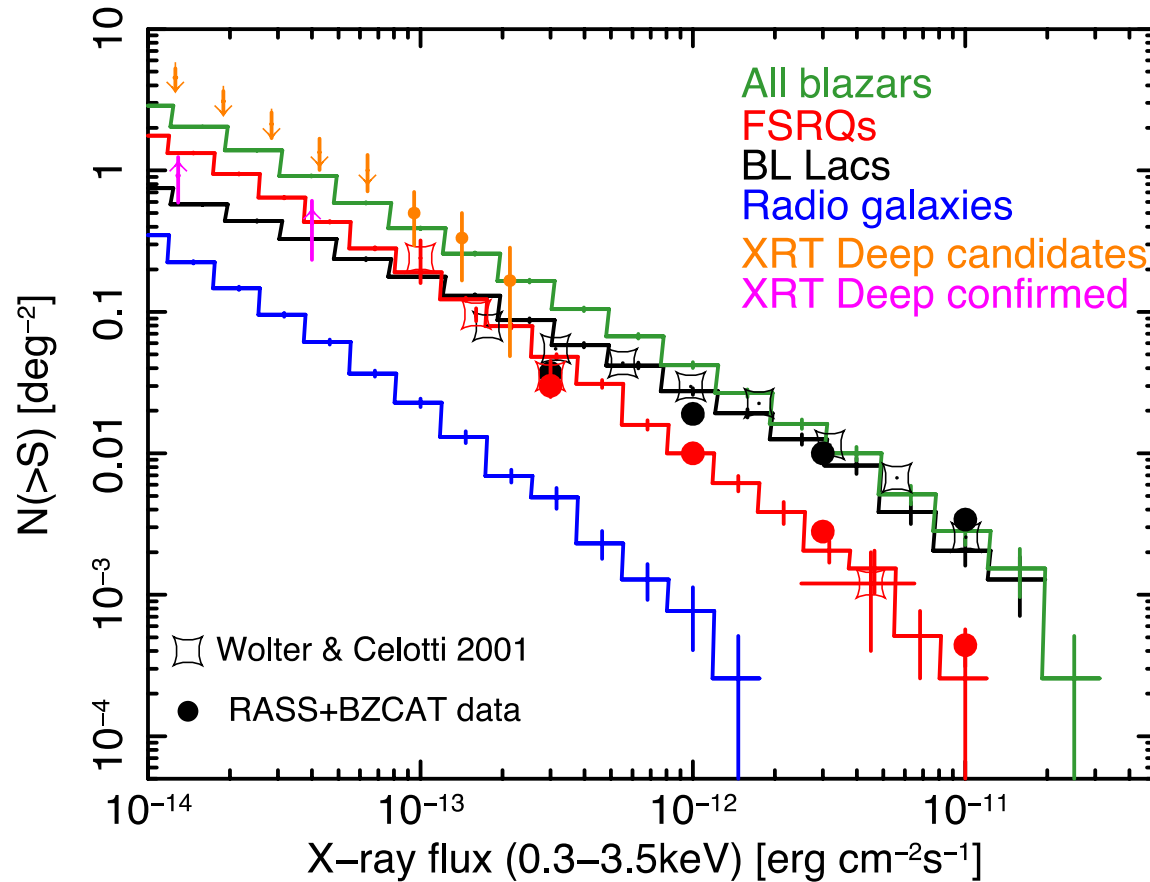


A.Read 25/08/17
Theseus #2 DOY:1-365

Baldi et al. 2001





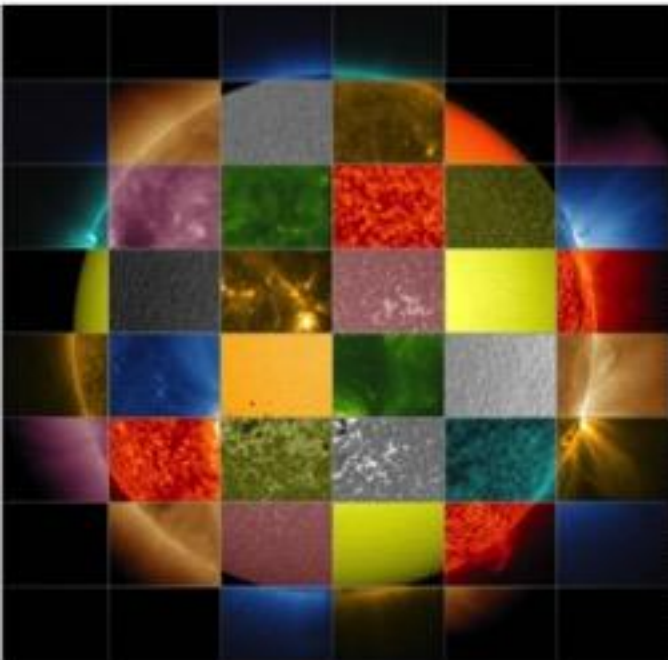




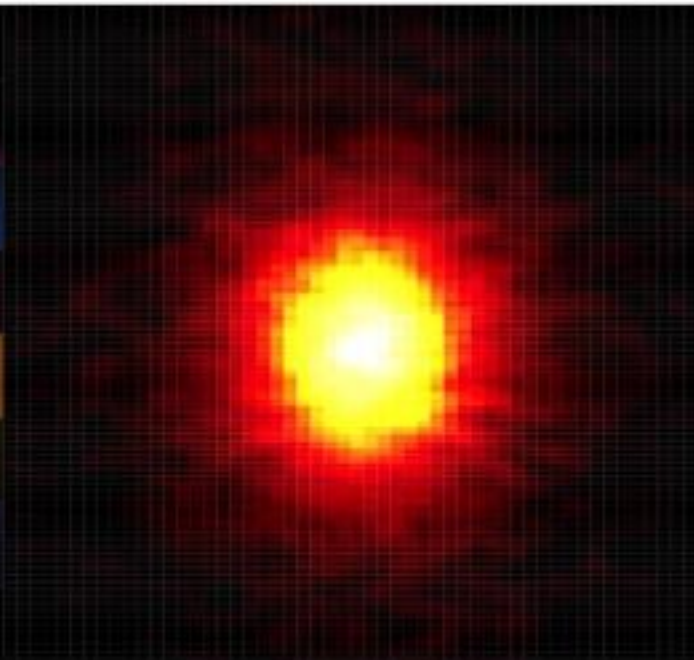
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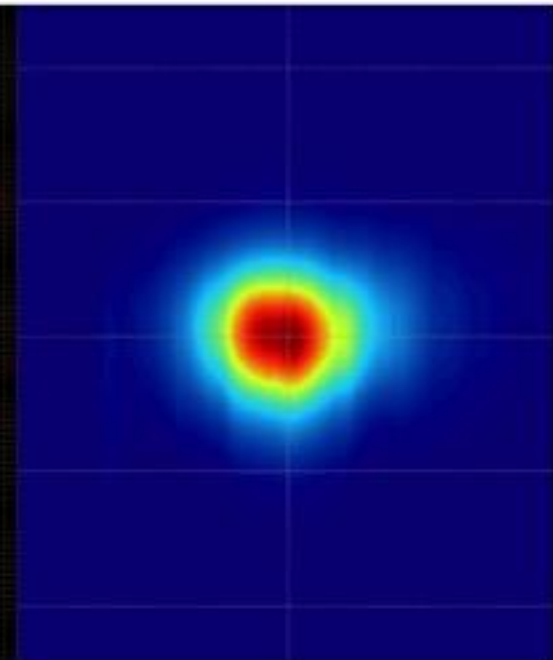
Multi-messenger pictures of our Sun



Multi-frequency
Electromagnetic radiation
NASA



Neutrinos
SuperKamiokande



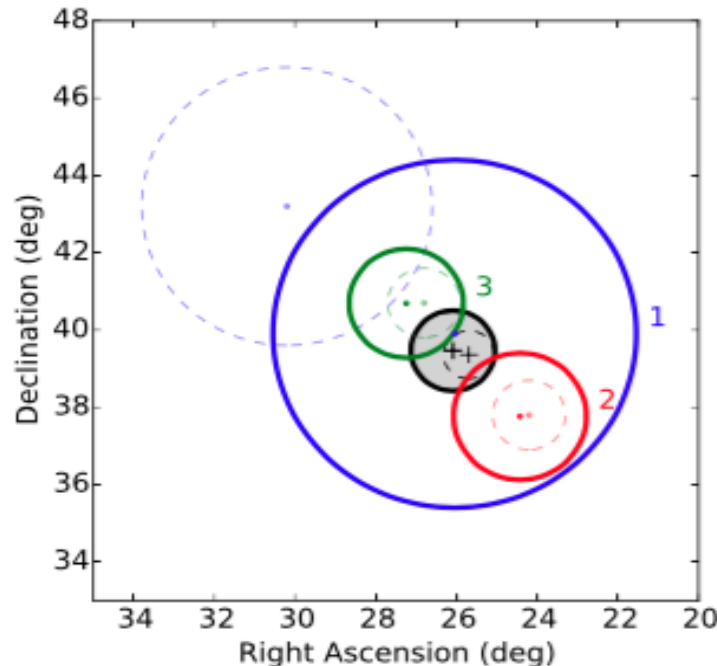
Cosmic Rays
Moon Mapping Project
Data from Chang'e-1

Multiwavelength follow-up of a rare IceCube neutrino multiplet

IceCube: [M. G. Aartsen^{\[2\]}](#), [M. Ackermann^{\[114\]}](#), [J. Adams^{\[28\]}](#), [J. A. Aguilar^{\[16\]}](#), [M. Ahlers^{\[66\]}](#), [M. Ahrens^{\[99\]}](#), [I. Al Samarai^{\[42\]}](#), [D. Altmann^{\[40\]}](#), [K. Andeen^{\[68\]}](#), ...

Aartsen et al. 2017, arXiv: 1702.0613

On February 17 2016, the IceCube real-time neutrino search identified, for the first time, three muon neutrino candidates arriving within 100 s of each other which are consistent with a point source origin. Such a triplet is expected only once every 13.7 years as a random coincidence of background events. However, considering the lifetime of the follow-up program the probability to detect at least one triplet from atmospheric backgrounds is 32%. Follow-up observatories were notified in order to search for an electromagnetic counterpart. Observations were obtained by



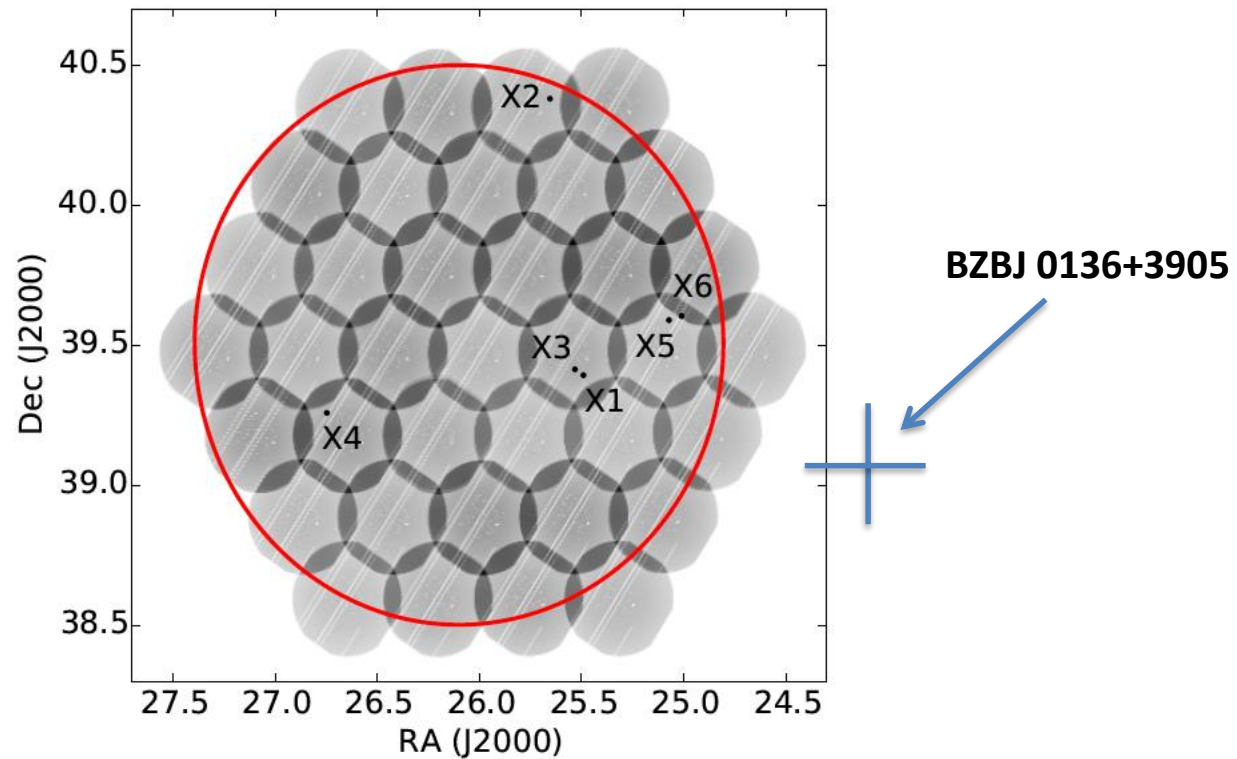


Fig. 3: Exposure map of the 37 *Swift* XRT pointings averaging 320 s per tiling. The red circle shows the 50% confidence bound to the triplet position. XRT sources (compare Table 3) are shown as black points.

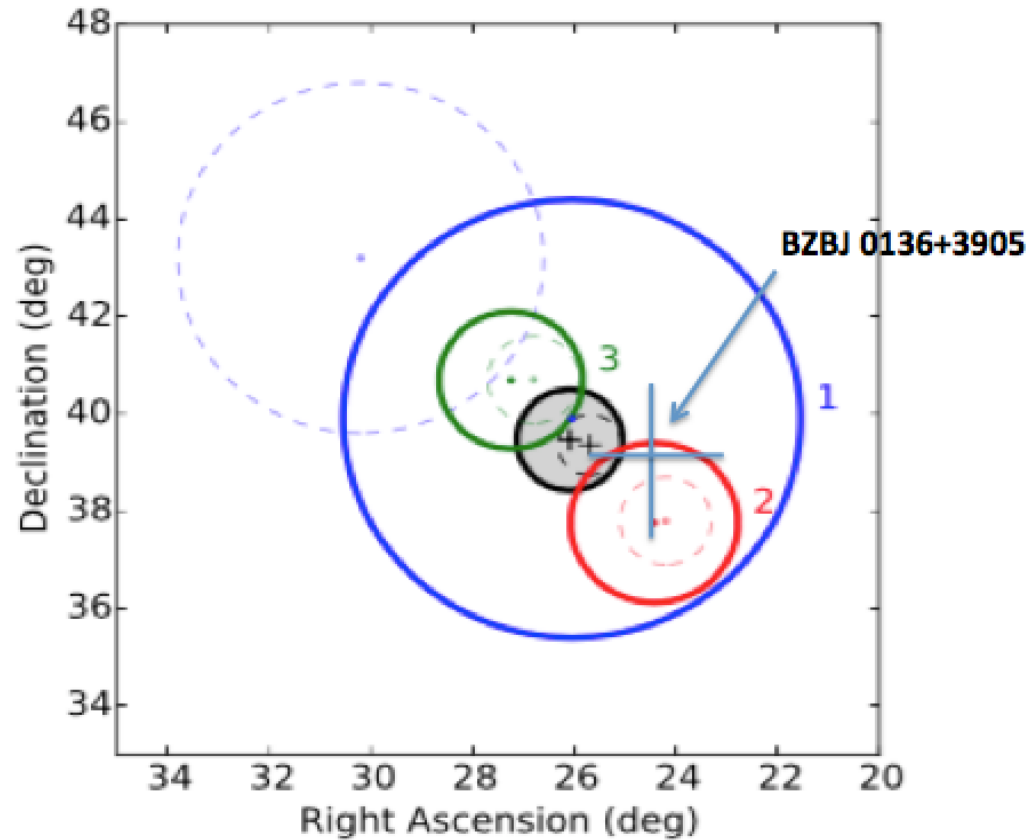
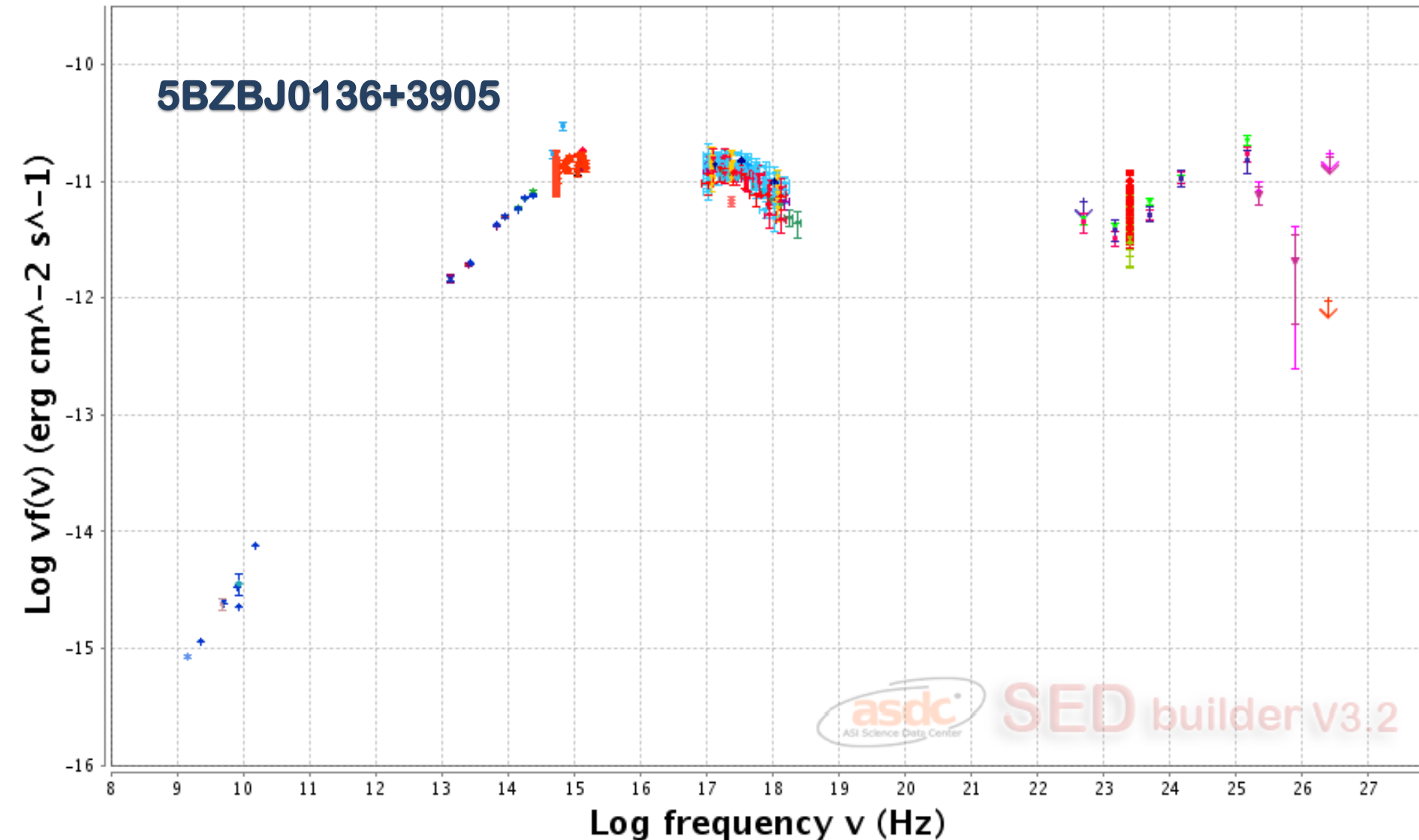
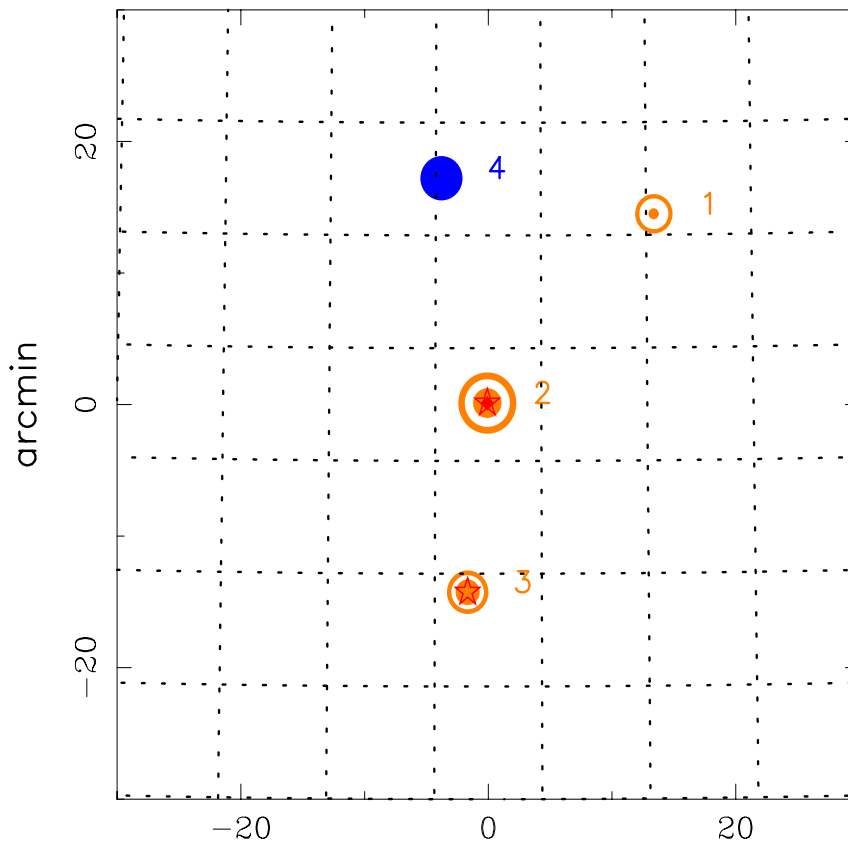
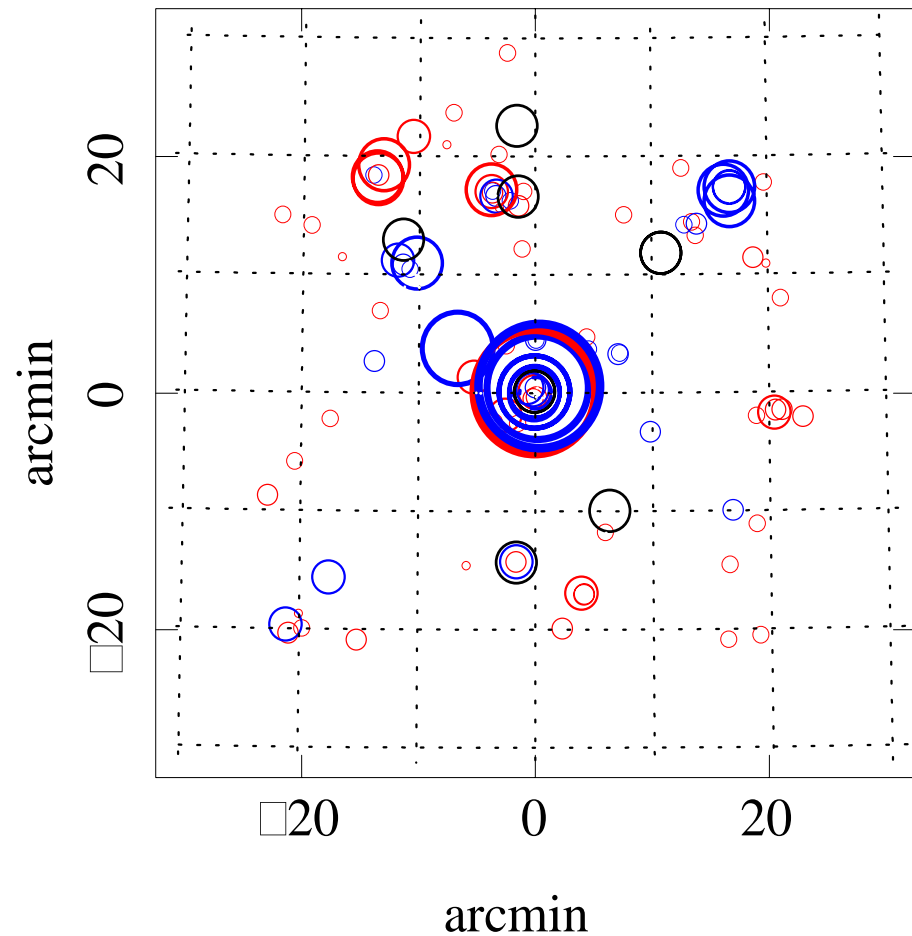


Fig. 1: Location of the three neutrino candidates in the triplet with their 50% error circles. The plus sign shows the combined direction and the shaded circle is the combined 50% error circle. The solid circles show the results of the MPE reconstruction which is as the default reconstruction in the following and the thin dashed circles correspond to the results of the Spline MPE reconstruction (compare Table [1](#)).



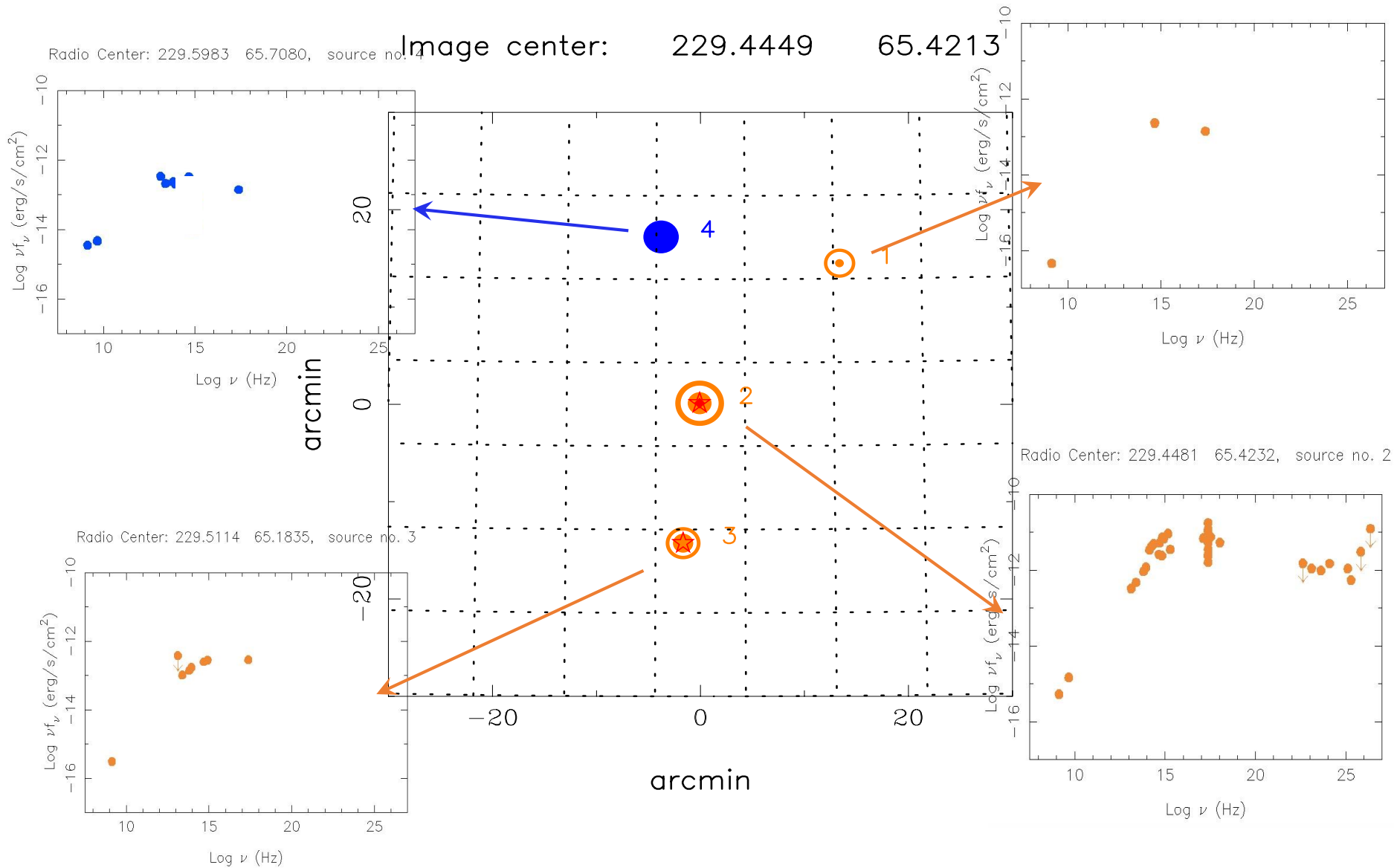
Fermi 3FHL J1517.7+6525



Radio Center: 228.9044 65.6620, source no. 1

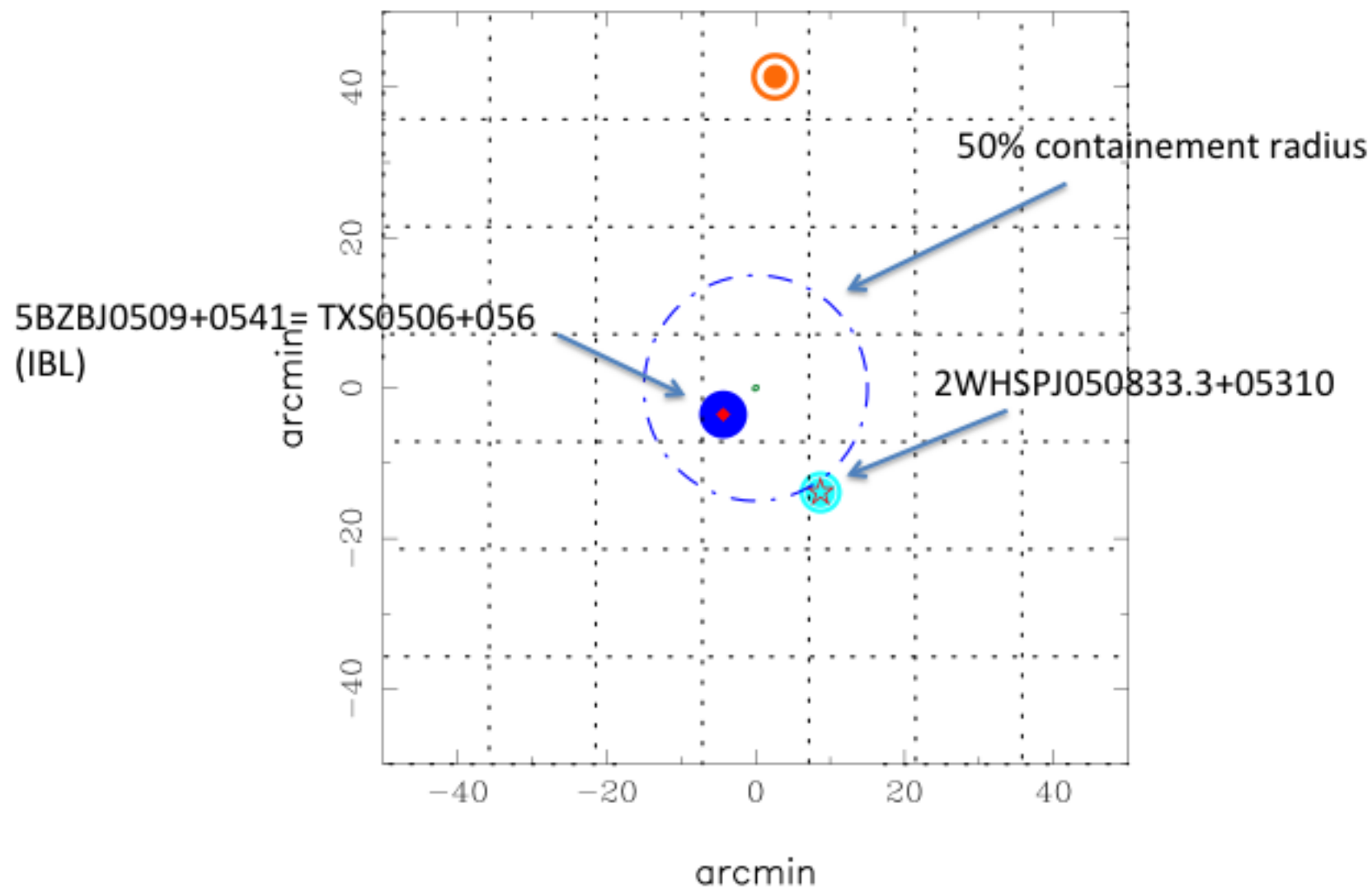
Radio Center: 229.5983 65.7080, source no. 4

Image center: 229.4449 65.4213

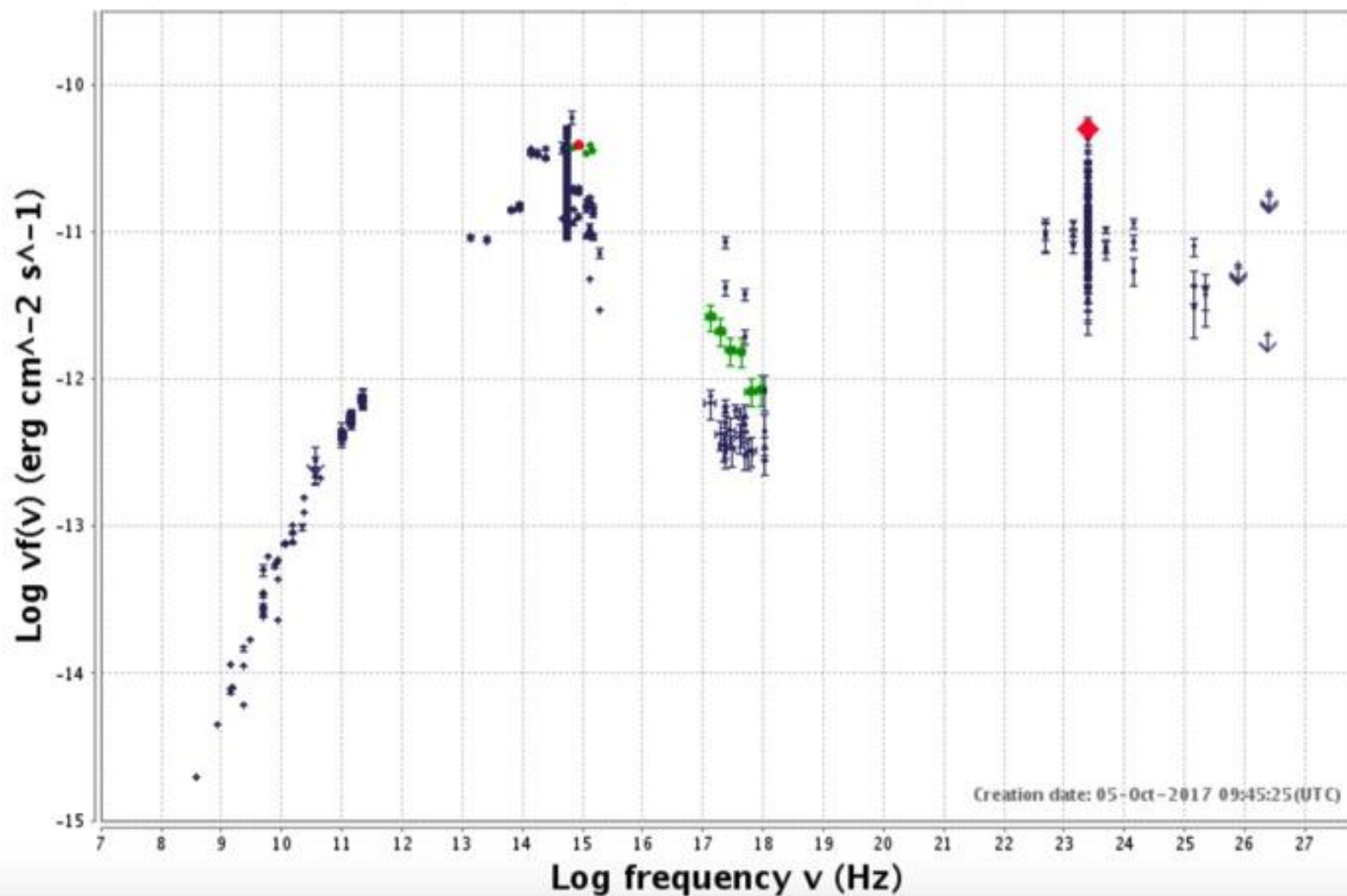


Neutrino ICE-Cube 170922

Image center: 77.2854 5.7517



5BZBJ0509+0541= TXS0506+056



Open UNiverse

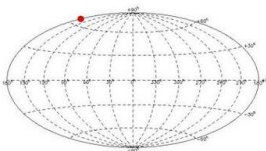
“Open Universe” is an initiative under the auspices of the UN COPUOS with the objective of stimulating a large increase of the utilization of space science data (e.g. astrophysics, planetary science, cosmic rays), extending the potential of scientific discovery to new participants in all parts of the world.

Open Universe was proposed by Italy at the 2016 COPUOS session where it was included among the activities in preparation of UNISPACE+50, in line with the thematic priority “Capacity Building”, with focus on Science, Technology, Engineering and Mathematics.

A very wide range of communities will benefit from Open Universe: professional scientists, citizen scientists, teachers and students, potentially any citizen interested in space science.

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Source Name(s) : **MKN421**
R.A.(J2000) = **11 04 27.34 (166.11392 deg)**
Dec.(J2000) = **+38 12 32.4 (38.209 deg)**

Prototype v.1.0

Object name or coordinates: MKN421 (ASDC)

MKN421

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Aladin Lite

Target:

166.11392 38.209

Surveys:

Fermi

GALEXGR6/AIS

DSS2

DSS2/red

DSS2/blue

SDSS9

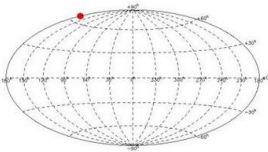
Mellinger

2MASS



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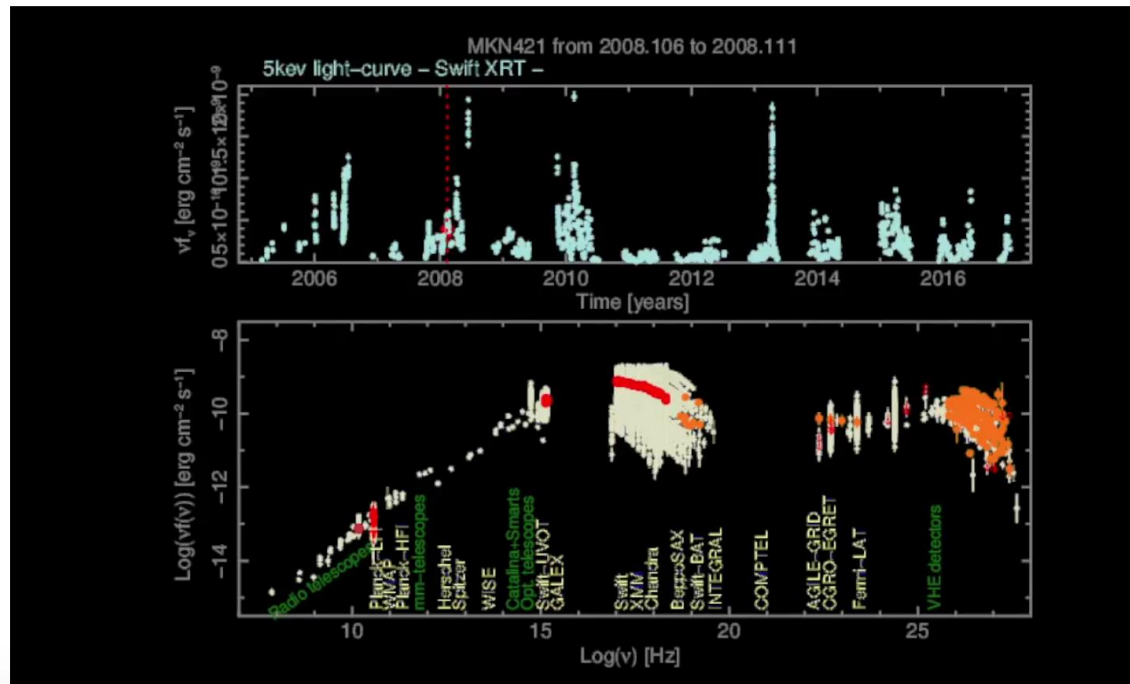
MKN421

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Conclusions

- Large all sky survey with $\sim 100,000$ - $300,000$ extragalactic sources
(Confusion limited?)
- $> 10,000$ blazars, a fair fraction of it potentially detectable by CTA
- Many flaring blazar to trigger IACTS (e.g. CTA) observations
- Blazar flaring data to compare with astrophysical neutrinos
- Plenty of variability studies of Jetted and non-jetted AGN