Piergiorgio Casella (INAF OA-Roma)

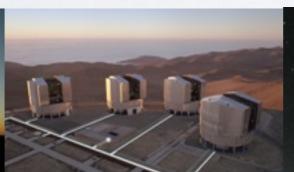
plan of the talk

- keep it short and simple
 - introduce LMXBs

- Low-Mass X-ray Binaries with
 - THESEUS

- what THESEUS can do
 - introduce SMARTNet



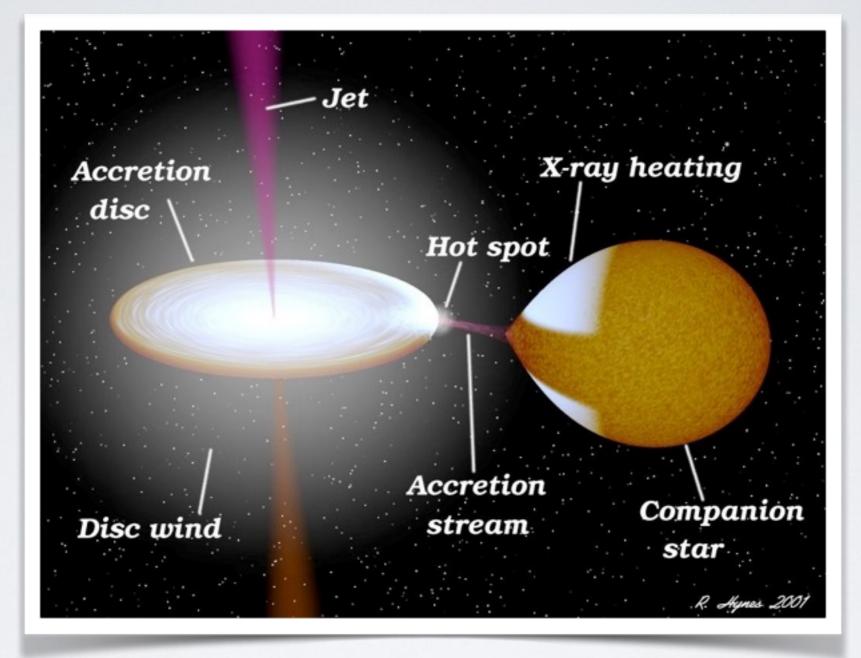








Low-Mass X-ray Binaries: Multi-Wavelength Sources

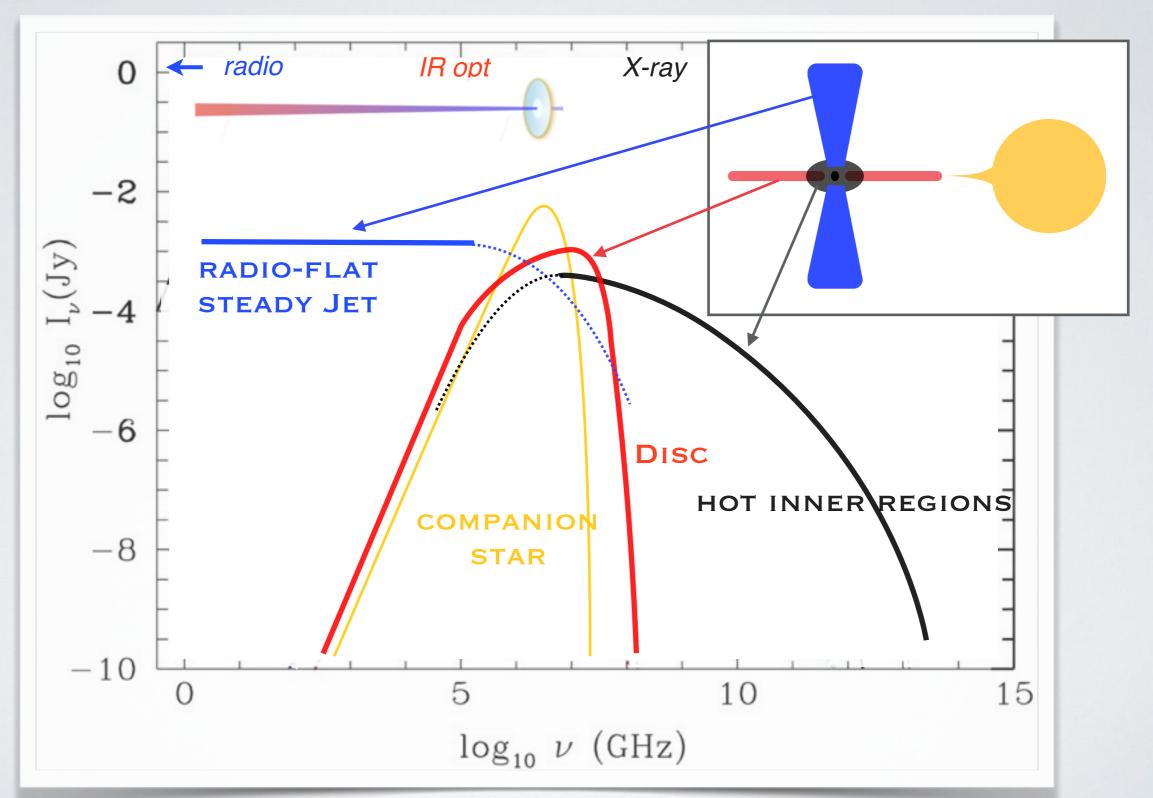


STUDY ACCRETION EVOLUTION ON PHD-TIMESCALES

COMPARISON WITH AGN / TDE

FEEDBACK ON ENVIRONMENT - STRONG GRAVITY

Low-Mass X-ray Binaries: Multi-Wavelength Sources

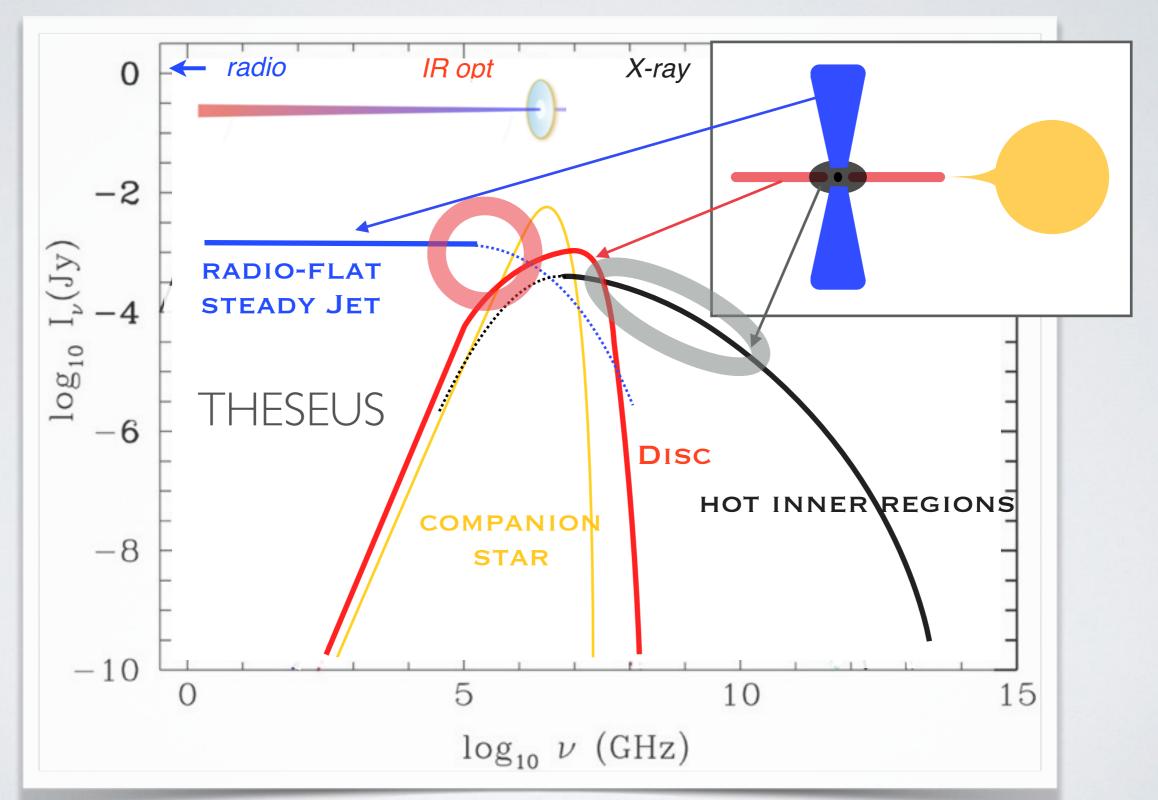








Low-Mass X-ray Binaries: Multi-Wavelength Sources



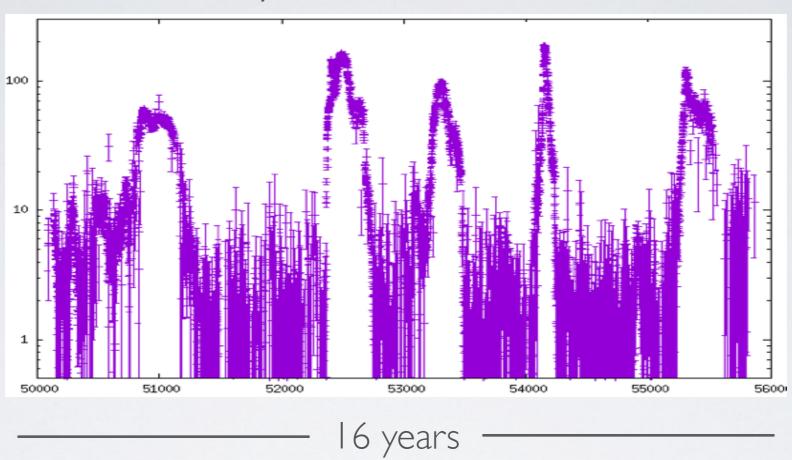






Low-Mass X-ray Binaries: Transient Sources

GX 339-4 RXTE/ASM 2-10keV



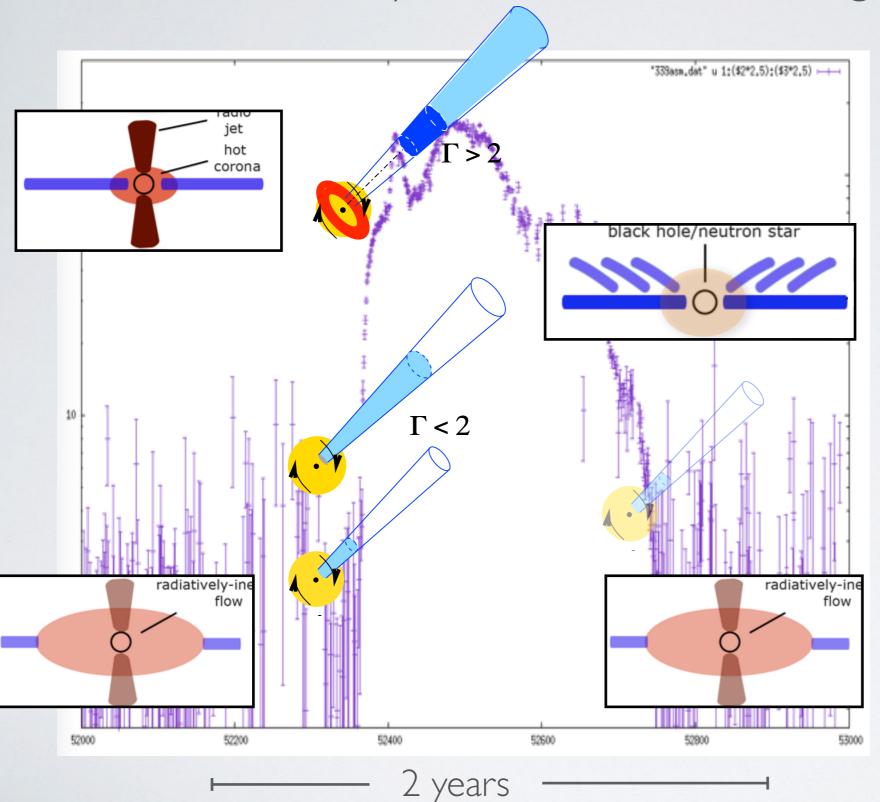
outburst rate: ranging from ~1/year to ~1/century outbursts duration: ranging from ~few weeks to ...persistent a few hundreds sources (~60 BH candidates + ~200 NS)

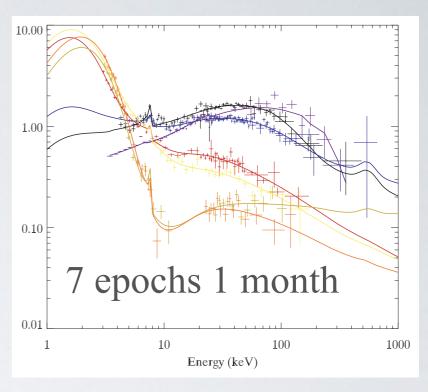


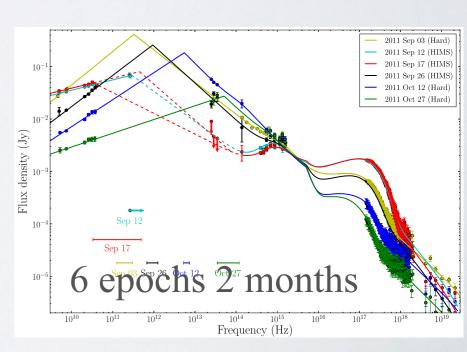




Low-Mass X-ray Binaries: Multi-Wavelength Variable Sources







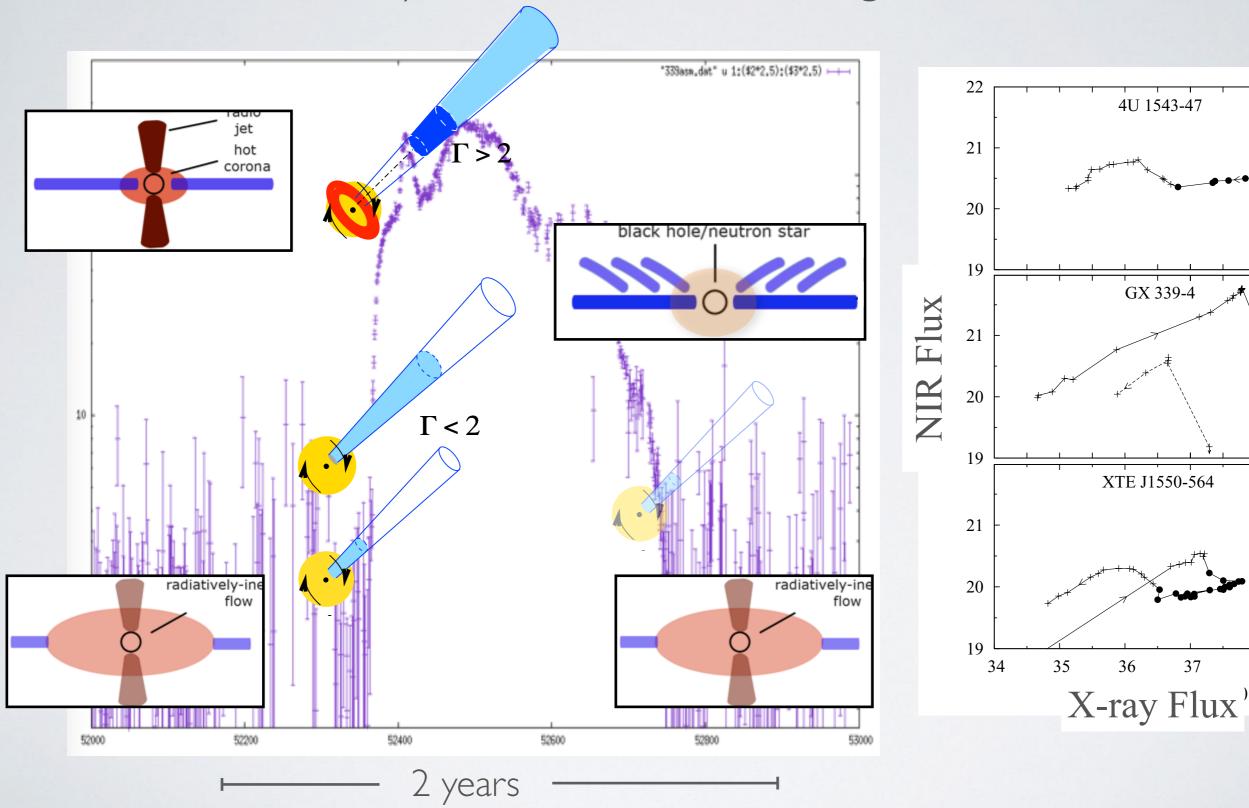


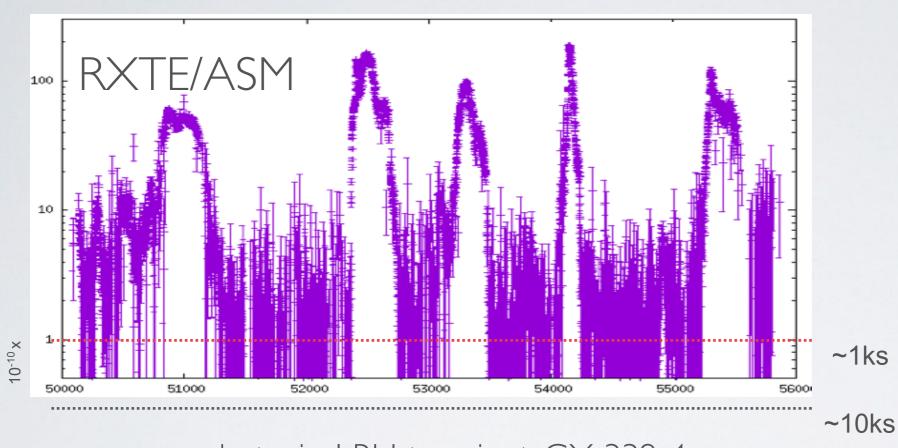




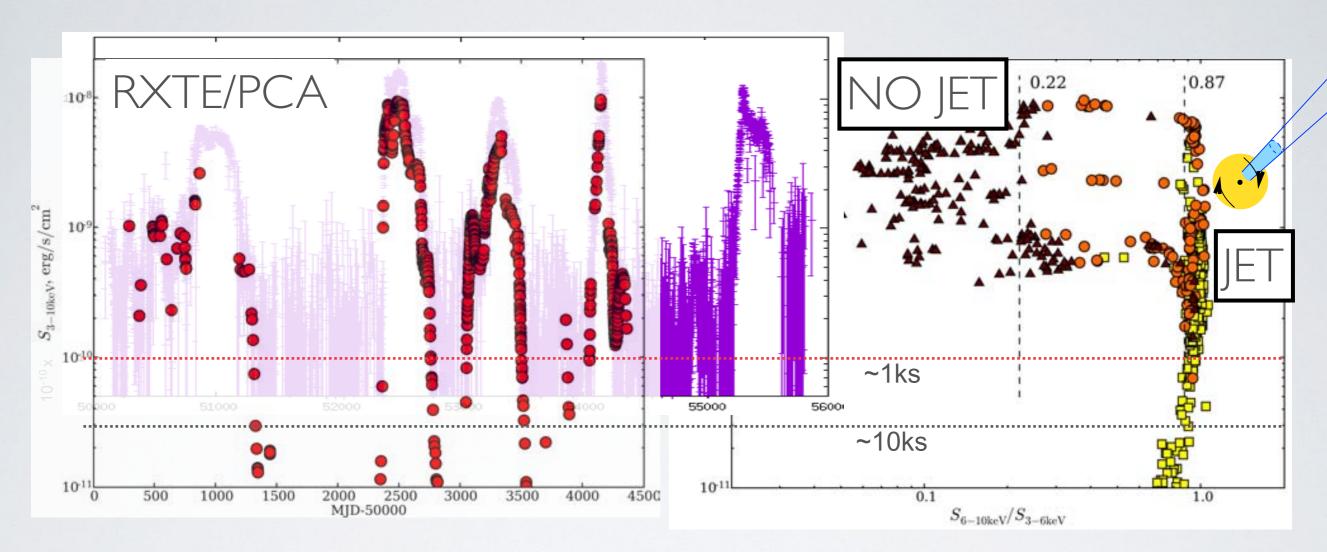
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Low-Mass X-ray Binaries: Multi-Wavelength Variable Sources





archetypical BH transient: GX 339-4



in several cases, IR flux always accessible with IRT...DOWNTO QUIESCENCE

THESEUS can cover the WHOLE outburst in both X-rays and IR

Key question I: WHERE are there?

- Triggers needed, for large-facility follow ups

Crucial point for XB community

we NEED to know when a XB goes on

we NEED to know what the source is doing, to trigger campaigns

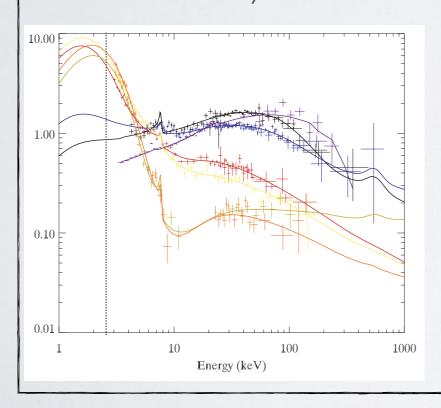
Key question I: WHERE are there?

Triggers needed, for large-facility follow ups

Key question 2: energy budget of accretion components

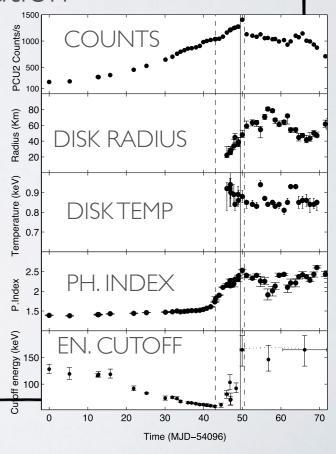
- X-ray sensitivities: XGIS can track the hot inflow evolution

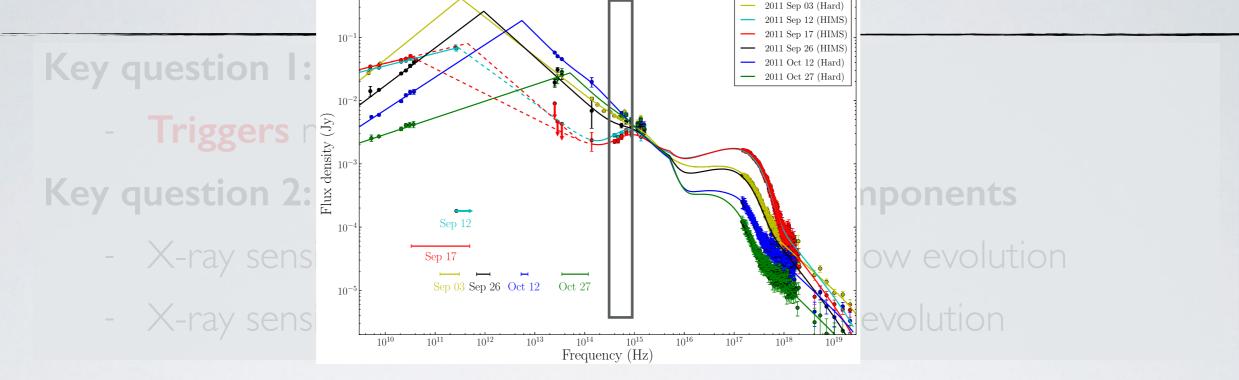
- X-ray sensitivities: **SXI** can track the soft disc evolution



THESEUS will do it for **all** (brights) **sources** over the full

useful energy band





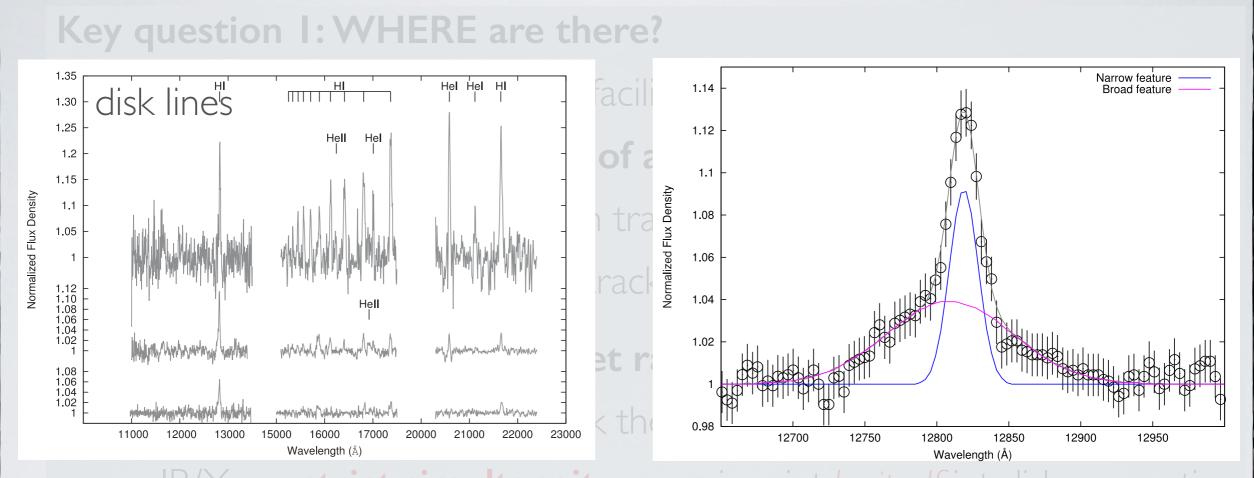
Key question 3: jet launching, jet radiative properties

- IR-sensitivities: IRT can track the jet spectral evolution
- IR/X-ray **strict simultaneity:** can pinpoint *by itself* jet-disk connection +SKA, LSST, etc.



+ 1 + + D





IR/X-ray strict simultaneity: can pinpoint by itself jet-disk connection

Bonus diagnostic: Disk IR lines! Wind IR lines?

+SKA, LSST, etc.

- IR-spectroscopy: bright NIR sources (H up to 14 or even 12)

Key question I: WHERE are there?

- Triggers needed, for large-facility follow ups

Key question 2: energy budget of accretion components

- X-ray sensitivities: XGIS can track the hot inflow evolution
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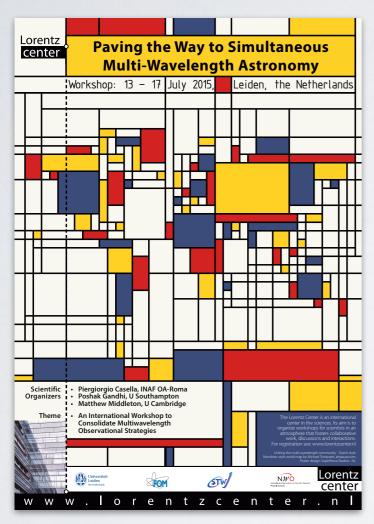
Bonus diagnostic: Disk IR lines! Wind IR lines? +SKA, LSST, etc.

- IR-spectroscopy: bright NIR sources (H up to 14 or even 12)

need tailored simulations, but prospects are very good

THESEUS in a Multi-Wavelength Context

REPORT/UPDATE ON A
COMMUNITY PROJECT
STARTED IN 2015



Coordination Team

Matthew Middleton (Southampton, UK)

Piergiorgio Casella (OAR, IT)

Poshak Gandhi (Southampton, UK)

Enrico Bozzo (ISDC, CH)

http://www.isdc.unige.ch/SMARTNet/

Simultaneous Multiwavelength Astronomy Research in Transients Network

A TOOL FOR THE COMMUNITY TO

- ★ optimise observing campaigns
- ★ speed up coordination for rapid transients
- **★** support collaborative approach
- ★ aid collecting & sharing information
- **★** boost communication in the community
- **★** develop community resources

What SMARTNet is

"Paving the way to simultaneous multi-wavelength astronomy"

New Astronomy Reviews

https://arxiv.org/abs/1709.03520

Everyone welcome to join (see website)

Google Search: smartnet astronomy

The more the merrier...

...it works if we are all there

What SMARTNet is NOT

NO binding policies, NO tight rules

- ★ use it only if you feel like it
- ★ share only what you want to share
- ★ handle the collaboration as you wish





